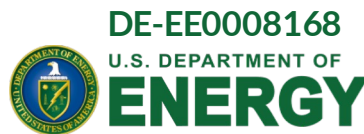


The influence of system-level design elements on convective cooling in PV solar farms



¹

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PV Performance Modeling and Monitoring Workshop
August 2022



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Portland State
UNIVERSITY



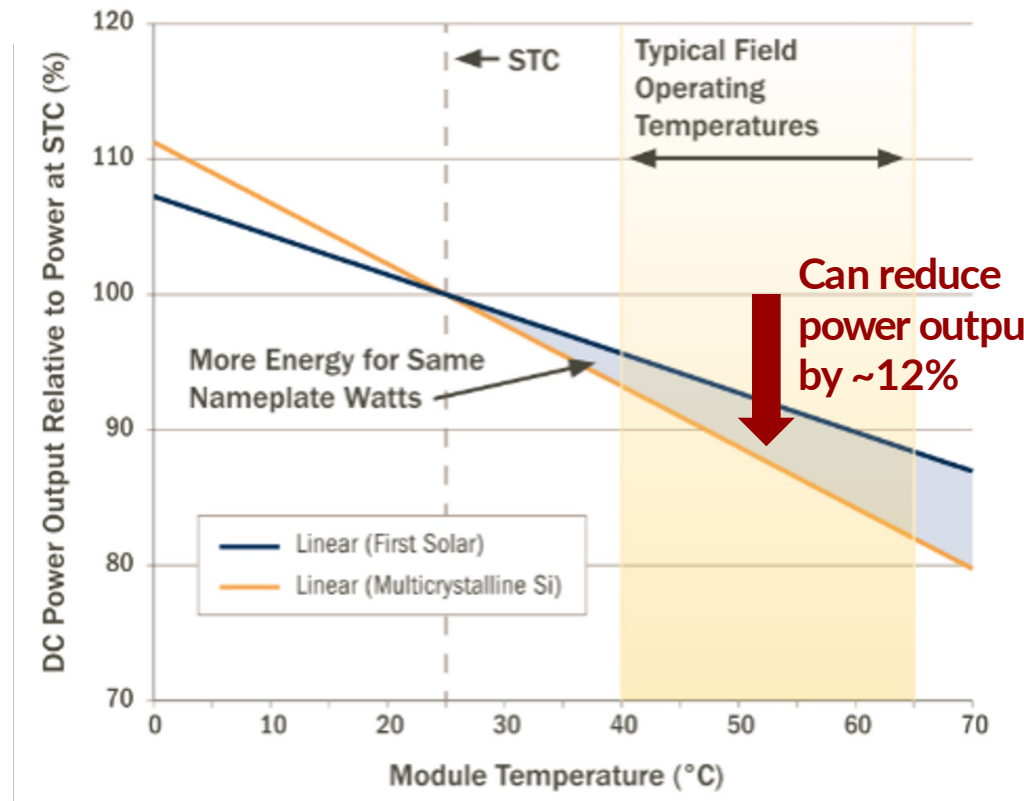
As PV module temperatures rise, efficiency drops and degradation accelerates

STC

1,000 W/m²

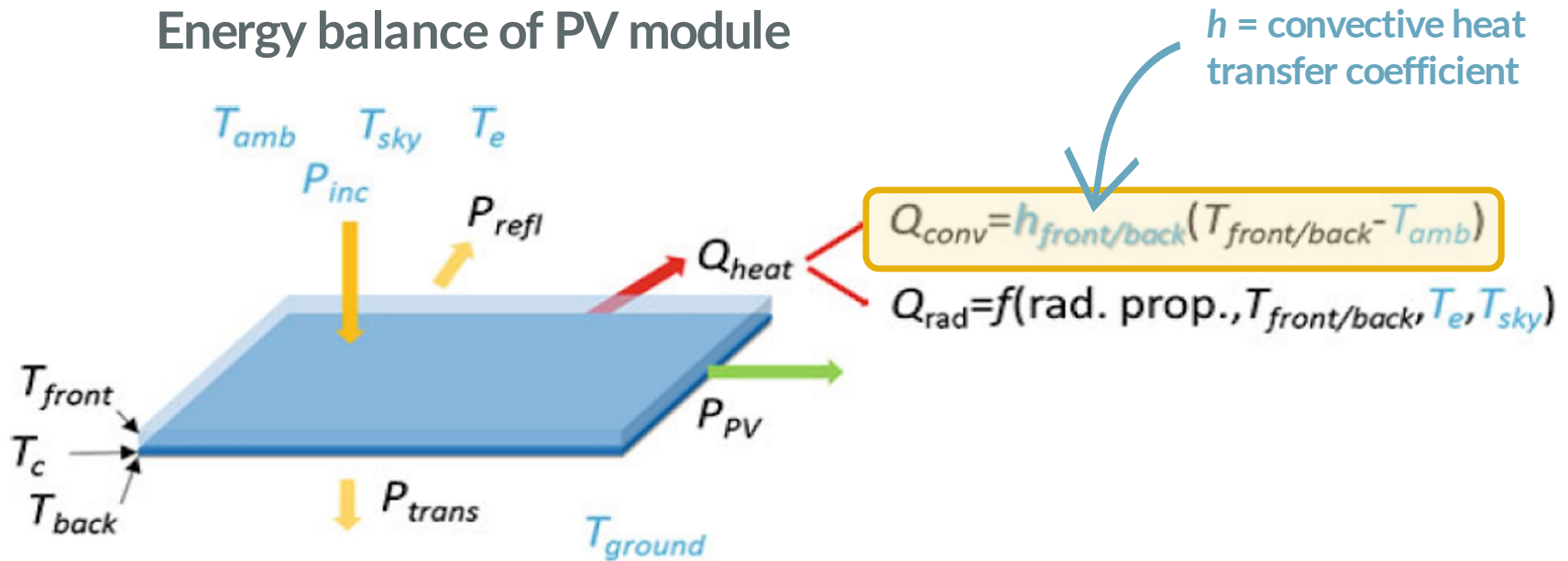
AM 1.5

25° C



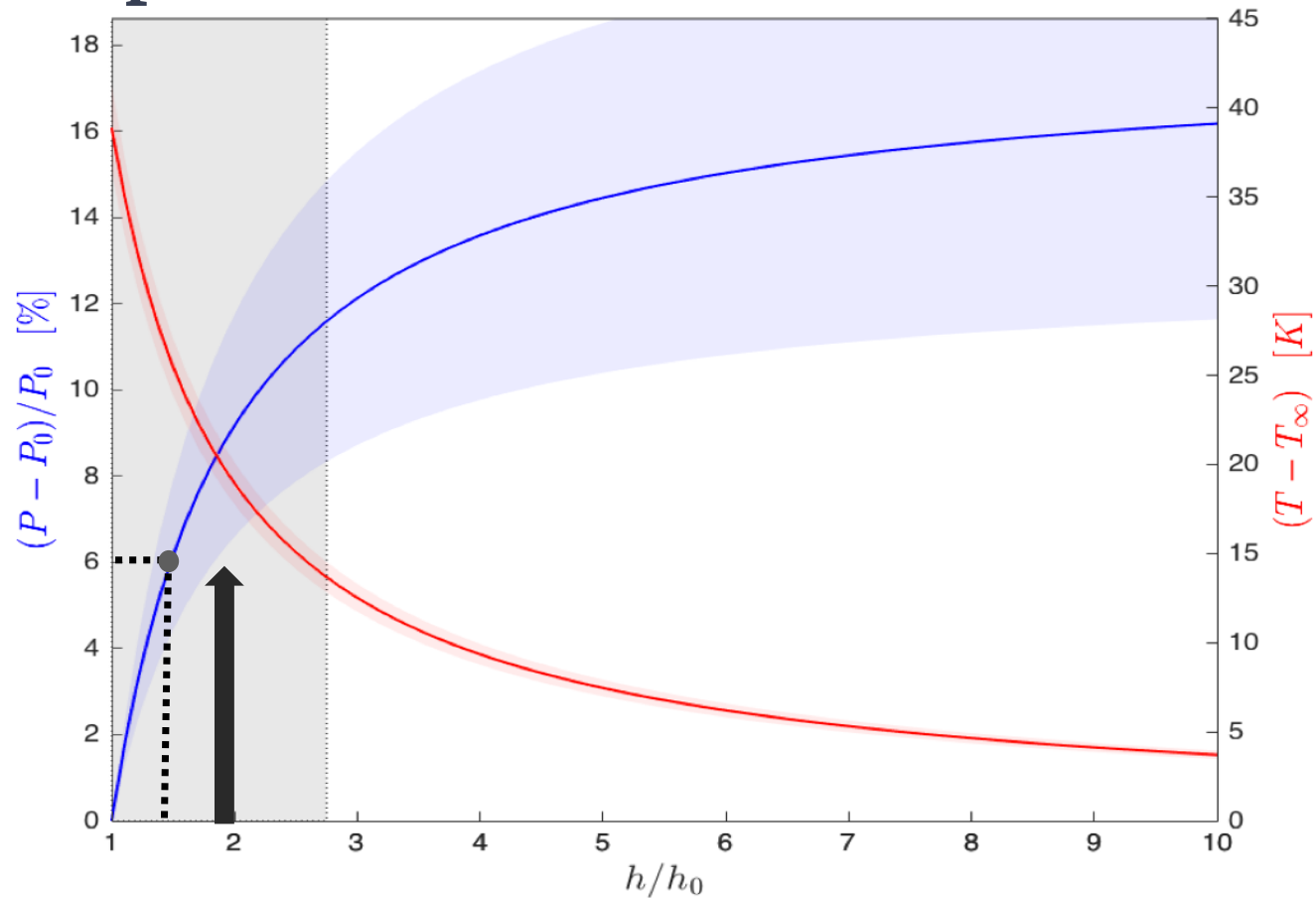
Dupré, O. (2016). *Physics of the thermal behavior of photovoltaic devices* (edited for clarity)

Enhancing convective heat transfer can cool solar modules



Dupre, Vaillon, *Thermal Behavior of Photovoltaic Devices*,
2017

Increasing convective heat transfer can increase power output



Increasing h by 50% can lead to an increase in power of 6%

Vaillon et al. 2018, Nature Scientific Reports

h = convective heat transfer coefficient

Convective heat transfer depends on several factors

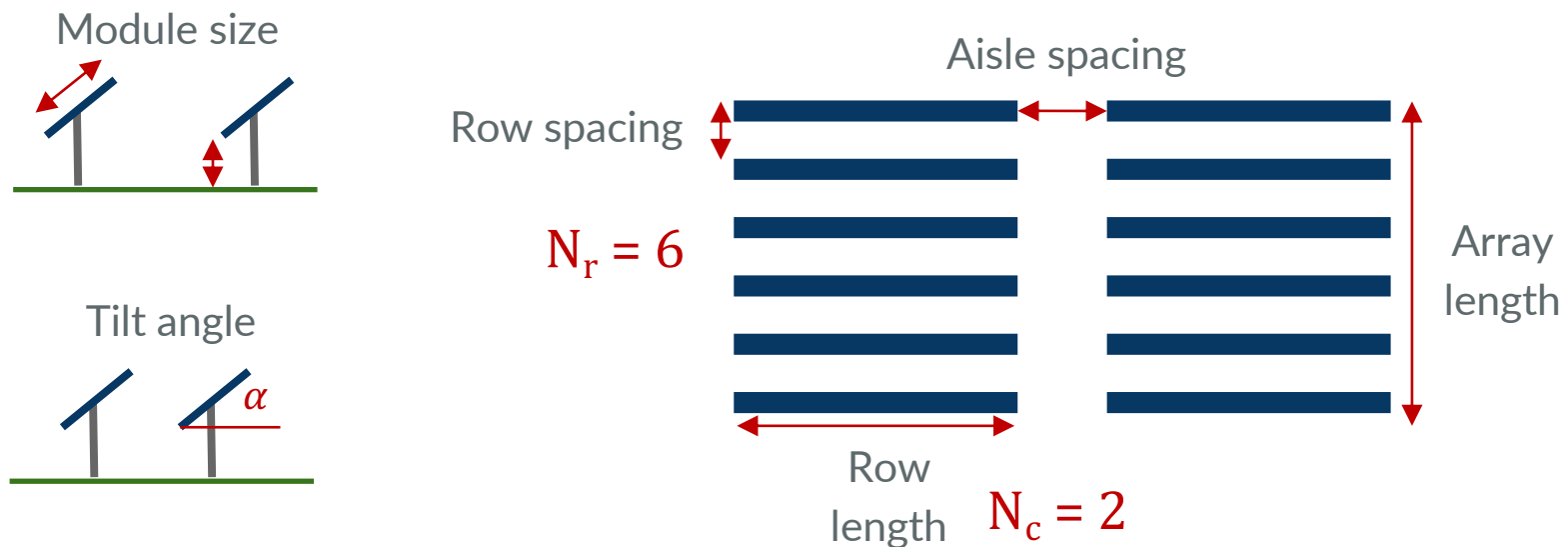
Environmental factors



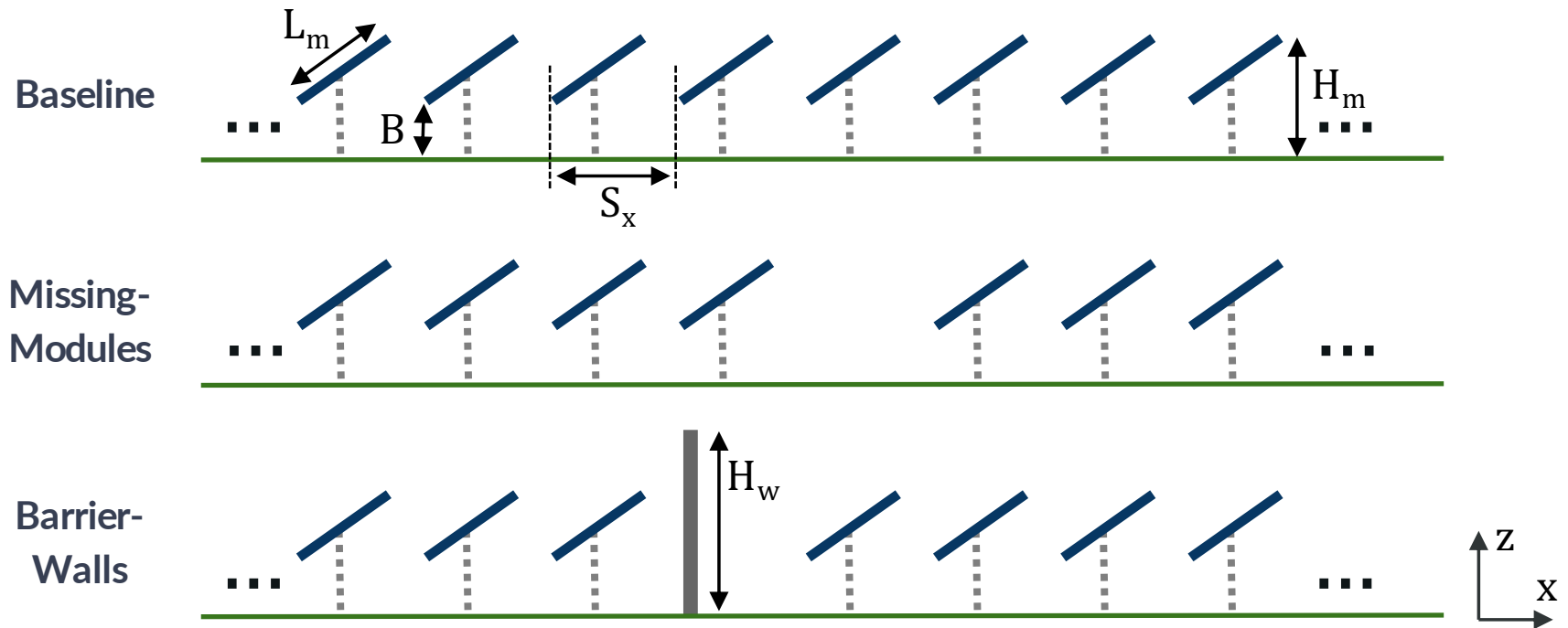
Geometric variables



Research Question: Can the arrangement of a solar farm naturally enhance convective heat transfer (h)?

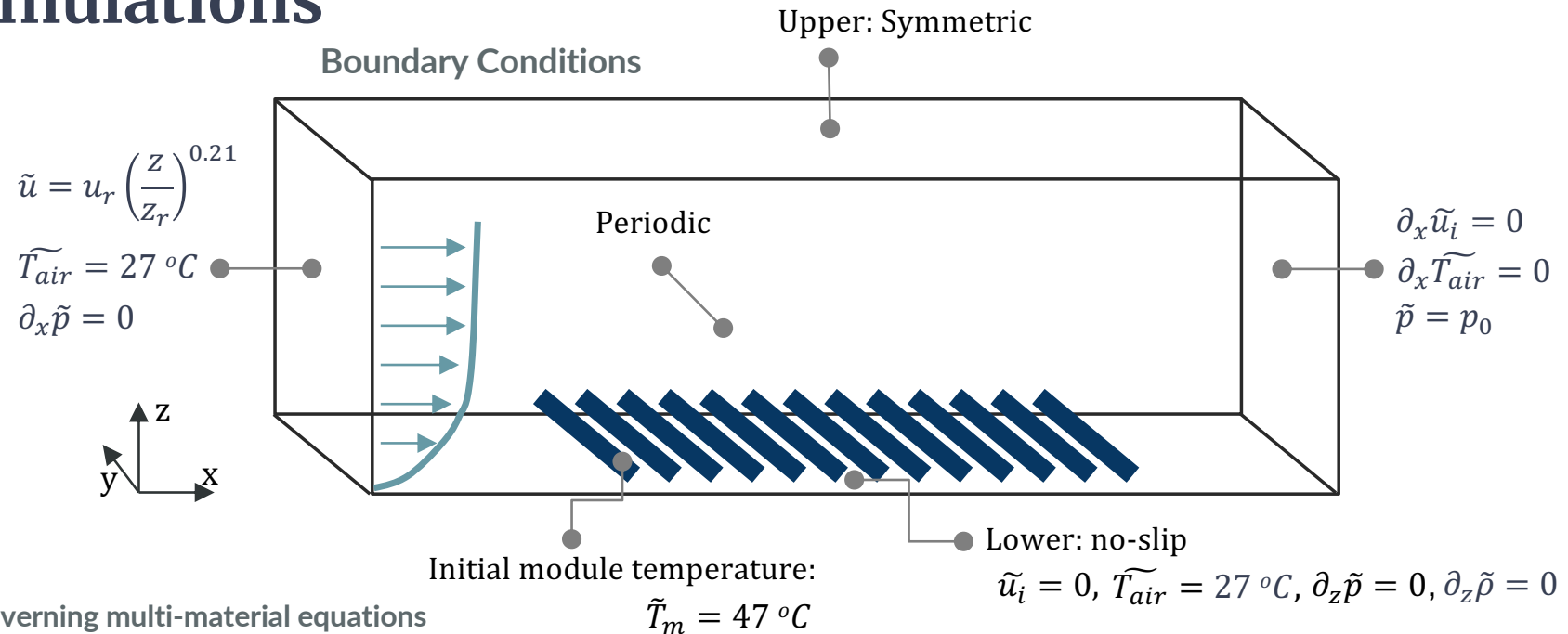


Three solar farm arrangements developed



Assembled module length: 3.3 m -- Tilt angle: 30 degrees -- Number of rows: 36

Uintah:MPMICE platform used for numerical simulations



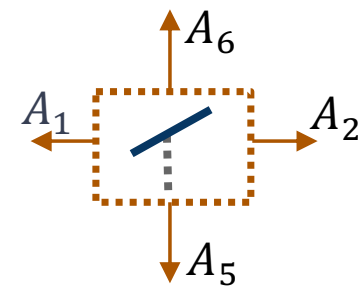
Governing multi-material equations

1. Conservation of mass
2. Conservation of momentum: Filtered Navier-Stokes equations in rotational form
3. Conservation of energy: temperature transport equation

Sub-grid model: Smagorinsky turbulence model

Inflow Turbulence: Generated with inflow grid turbulence generator (Hayati et al. 2017)

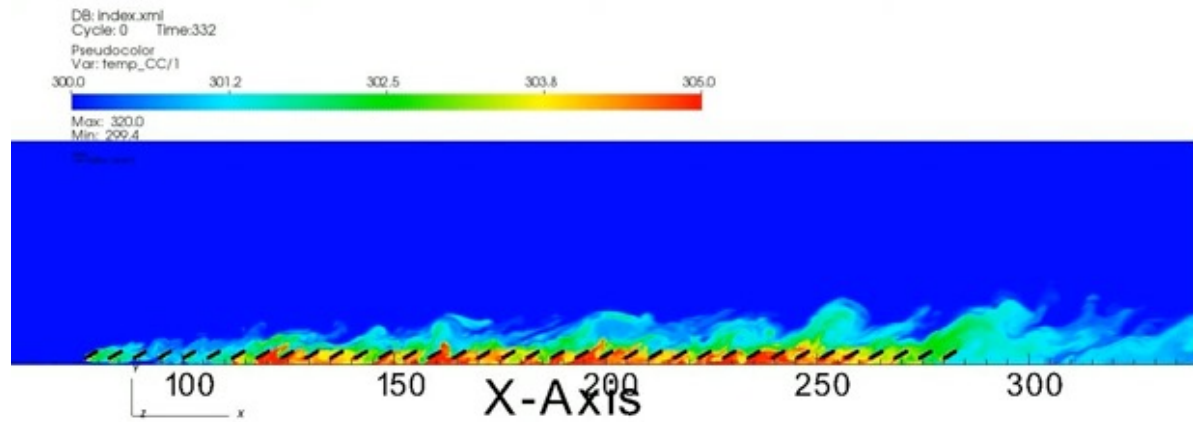
Flow solver: MPMICE (Material Point Method, Implicit, Continuous fluid, Eulerian) method: cell-centered, finite-volume, multi-material (Kashiwa & Rauen Zahn, LANL 1994, Sulsky et al. 1994 & 1995, Guilkey et al. 2007)



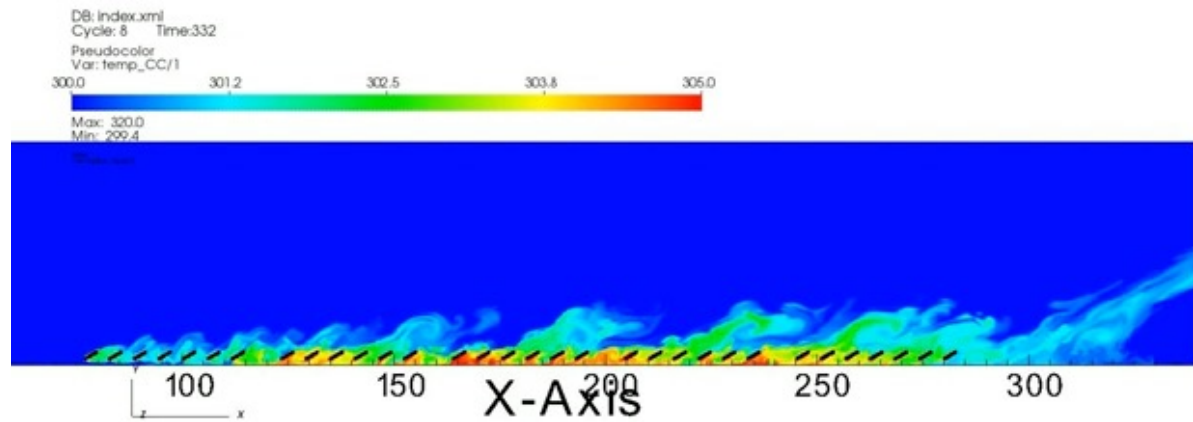
Newton's Law of Cooling

$$h = \frac{q_m}{T_m - T_\infty}$$

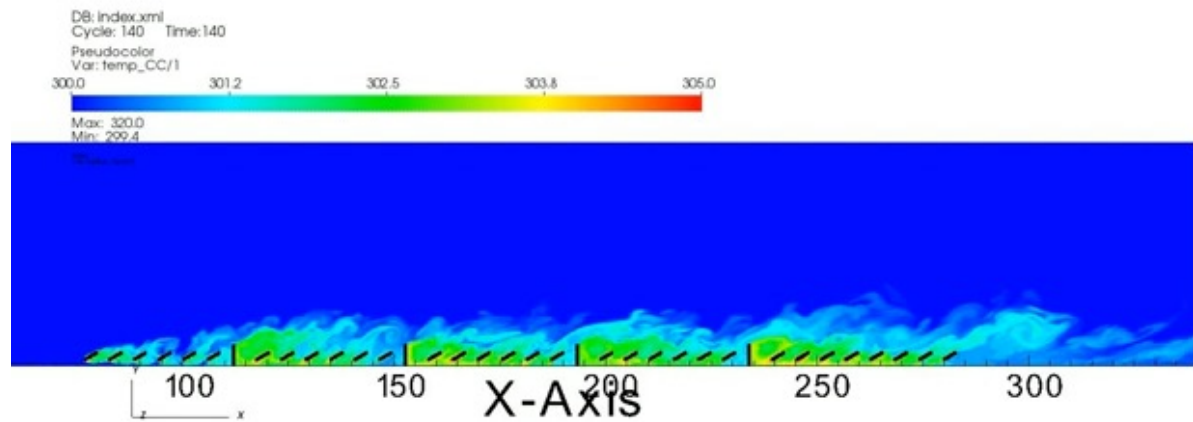
Baseline



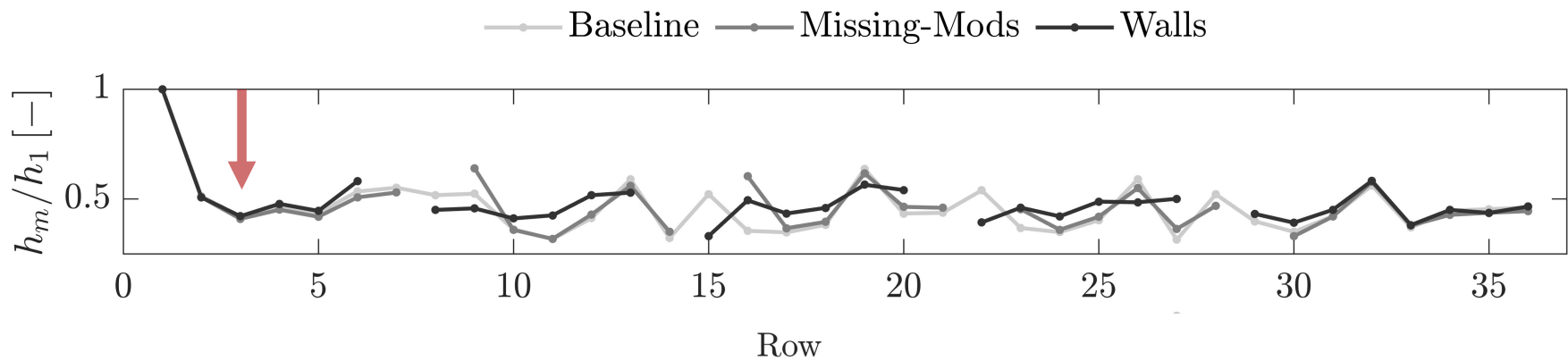
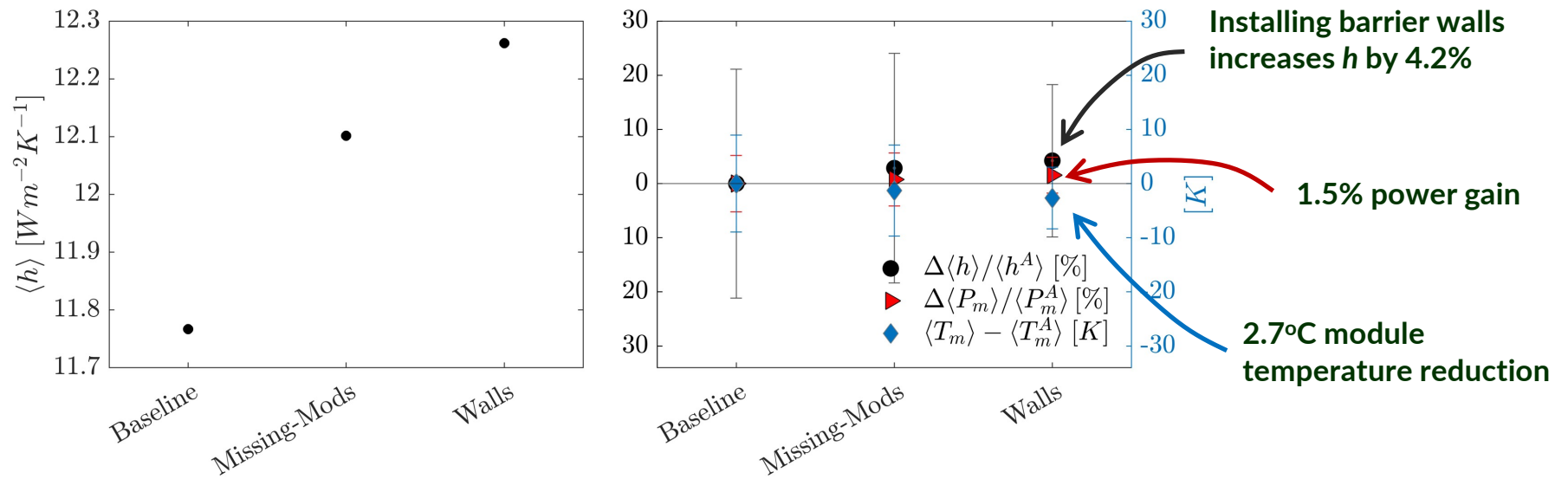
Missing-
Modules



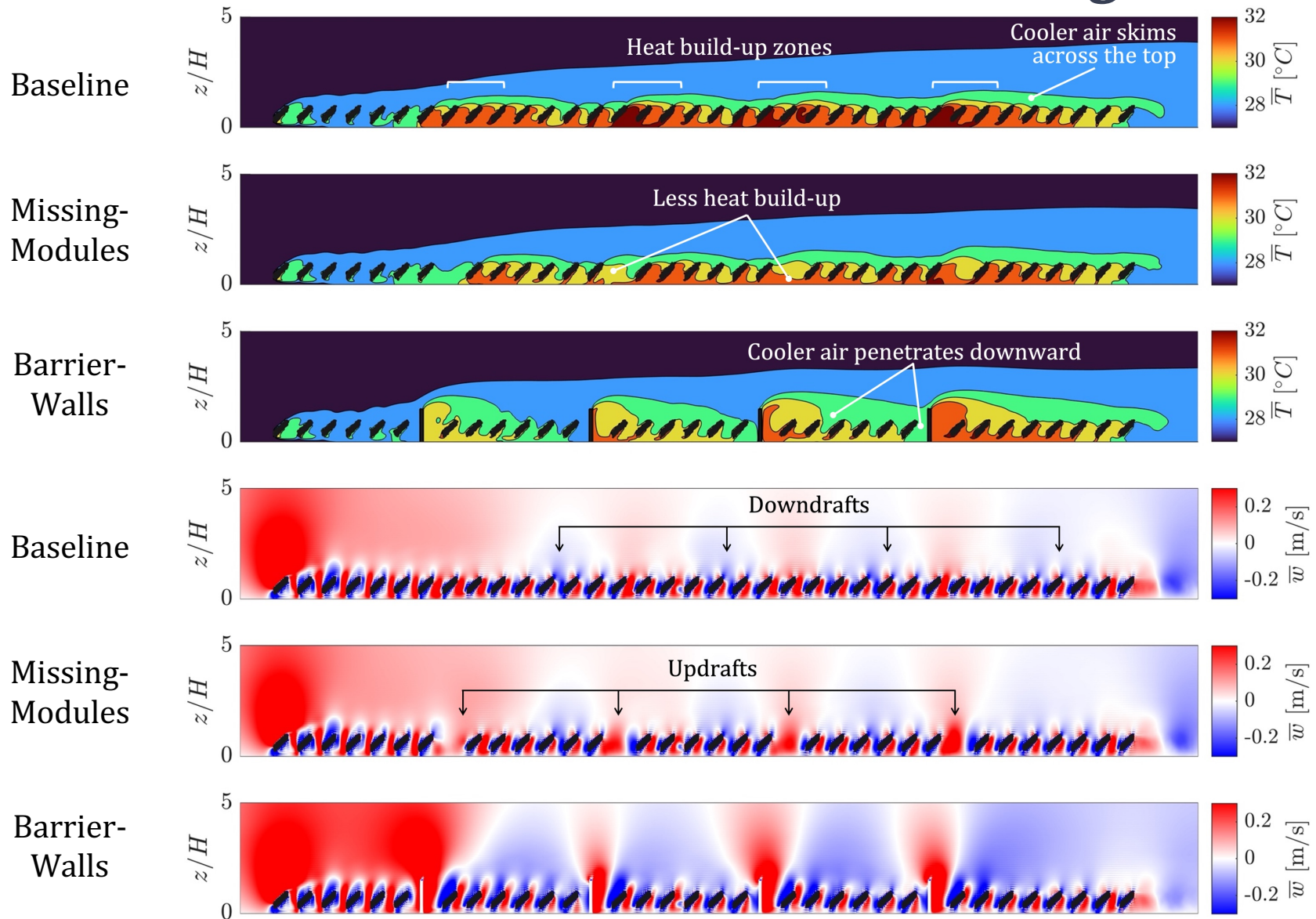
Barrier-
Walls



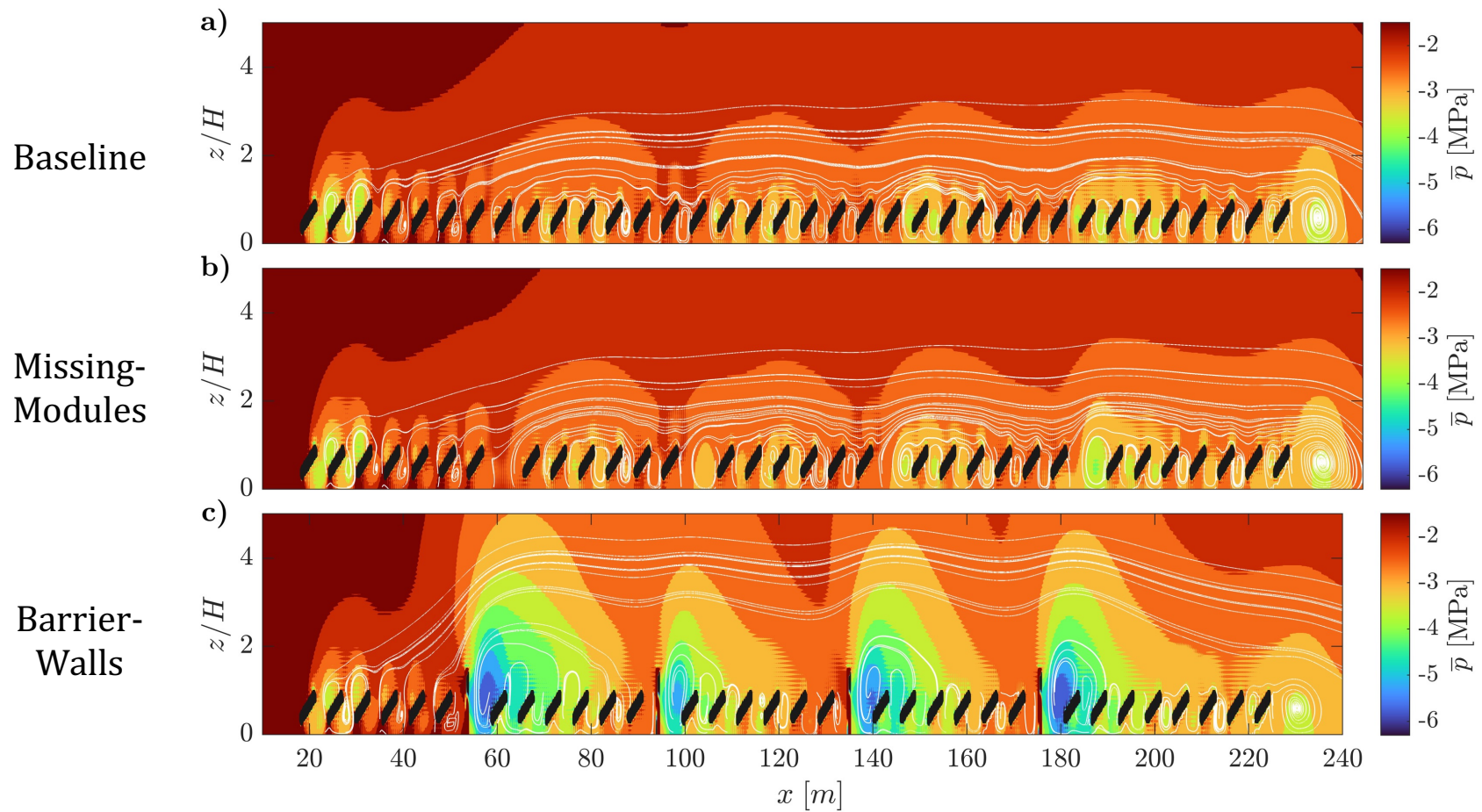
Barrier walls yield the greatest improvements



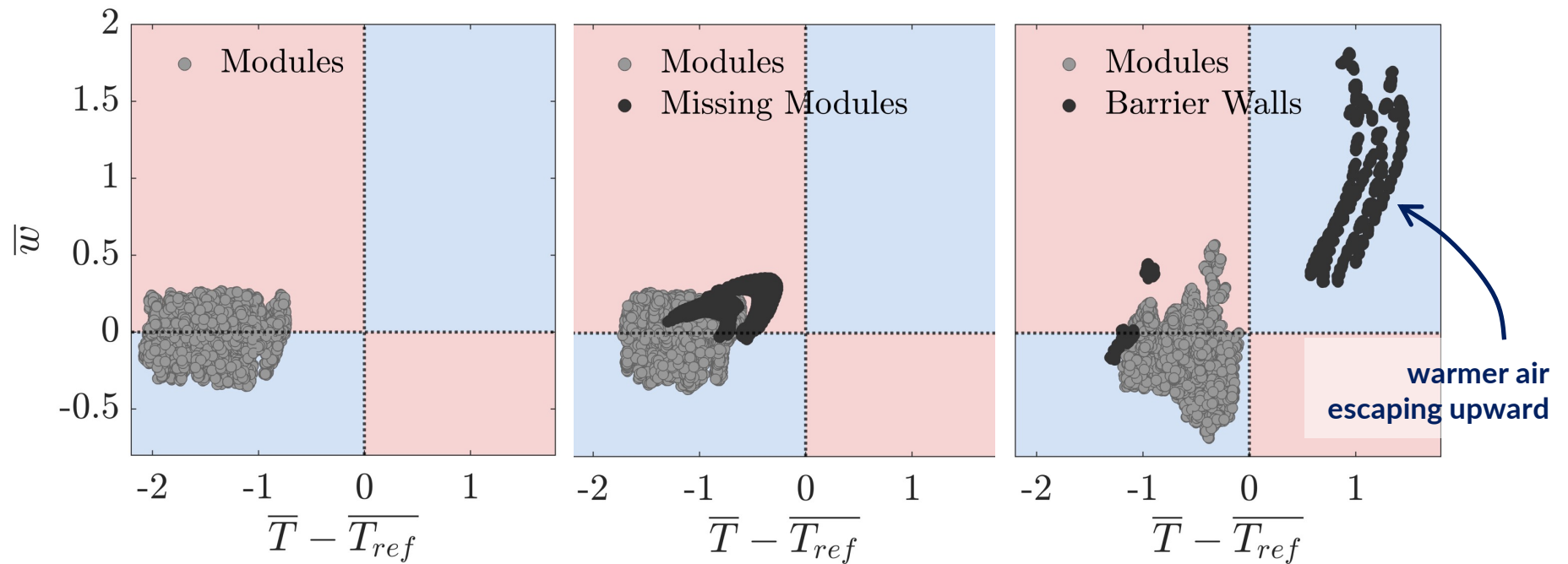
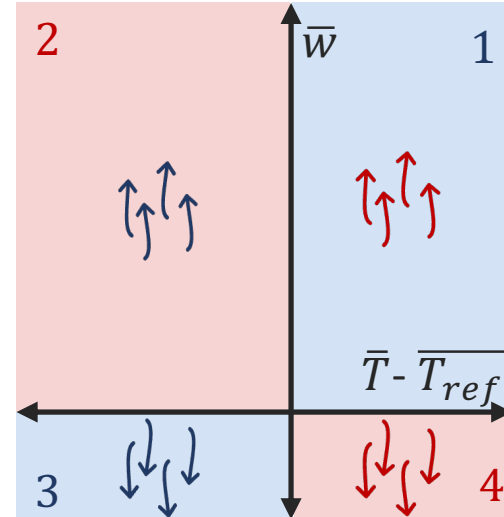
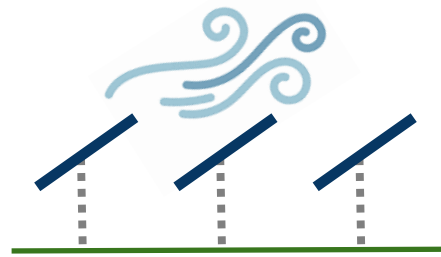
Thermal and flow behavior of each arrangement



Pressure fields that induce flow behaviors

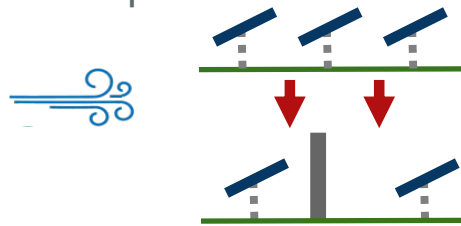


What flow behavior causes the enhanced cooling?

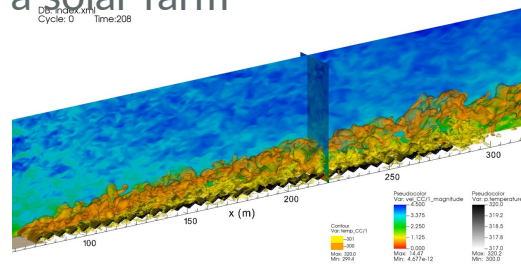


Contributions of this work

Low-cost cooling solution for PV plants



Spatial variations within a solar farm

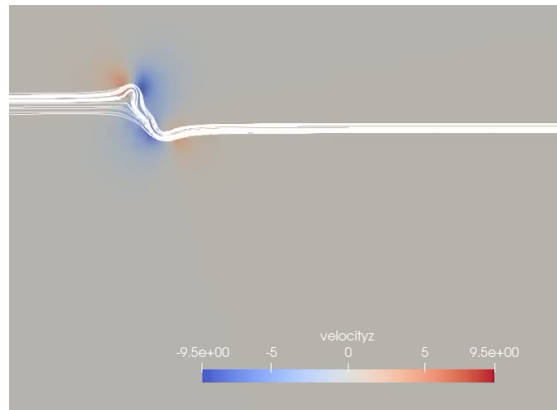


Related work

Wind loading of heliostats in concentrated solar power plants



July 26, 2017 - Heliostats at the Ivanpah Solar Project, owned by NRG Energy, Bright Source Energy, Bechtel and Google. (Photo by Dennis Schroeder / NREL)



Flow past a heliostat. (Video by Eliot Quon and Shashank Yellapantula / NREL 2022)

Check out these posters to learn more about this project

- 6) Prilliman, Matthew. "Technoeconomic Analysis of changing PV System Layout and Convection Heat Transfer"
- 7) Davis, Jace. "Convective Cooling of Solar Photovoltaic Modules in Unperturbed Atmospheric Conditions"
- 8) Glick, Andrew. "Effects of module configuration on convective cooling for utility-scale solar PV plants"

Thank you

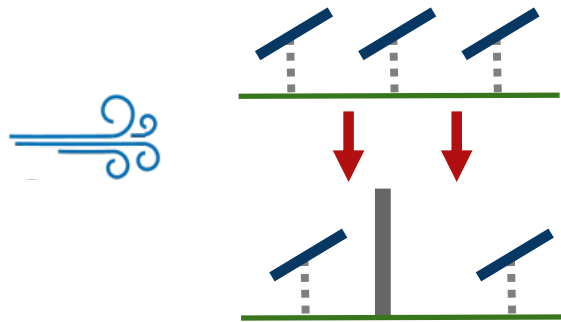
Brooke Stanislawski

brooke.stanislawski@nrel.gov

Back-Up Slides

Contributions of this work

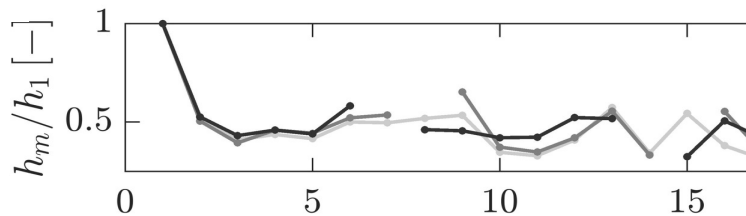
Low-cost cooling solution for PV plants



Manipulation of flow to induce cooling



Spatial variations within a solar farm



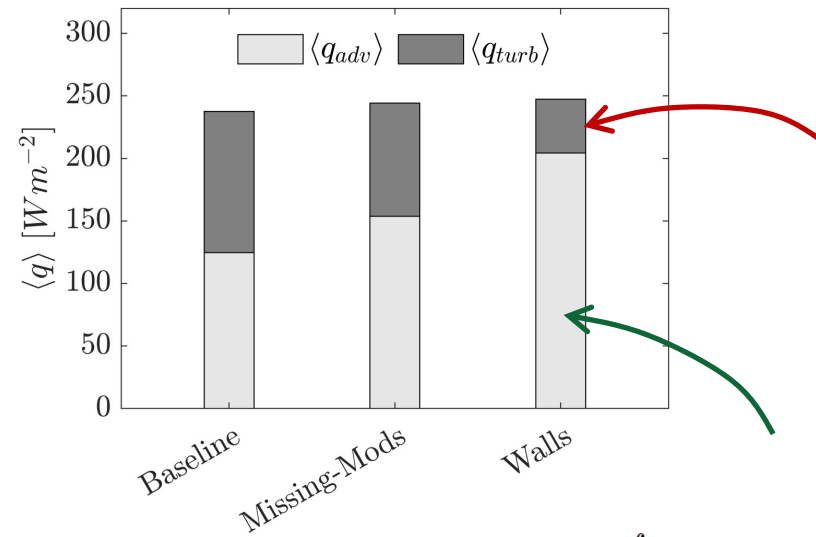
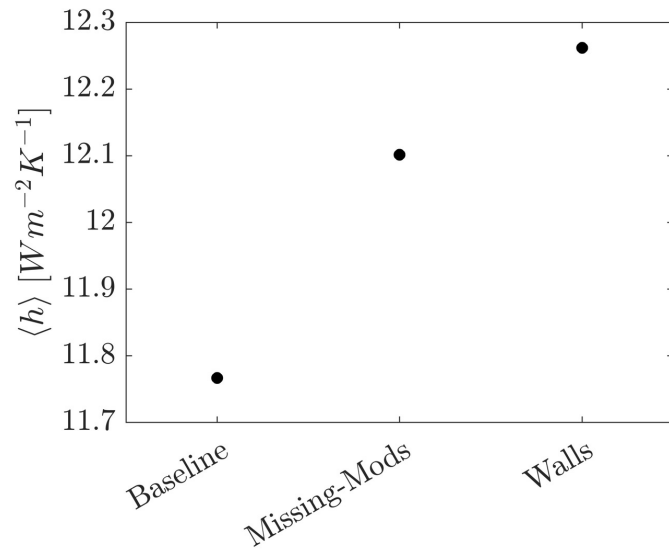
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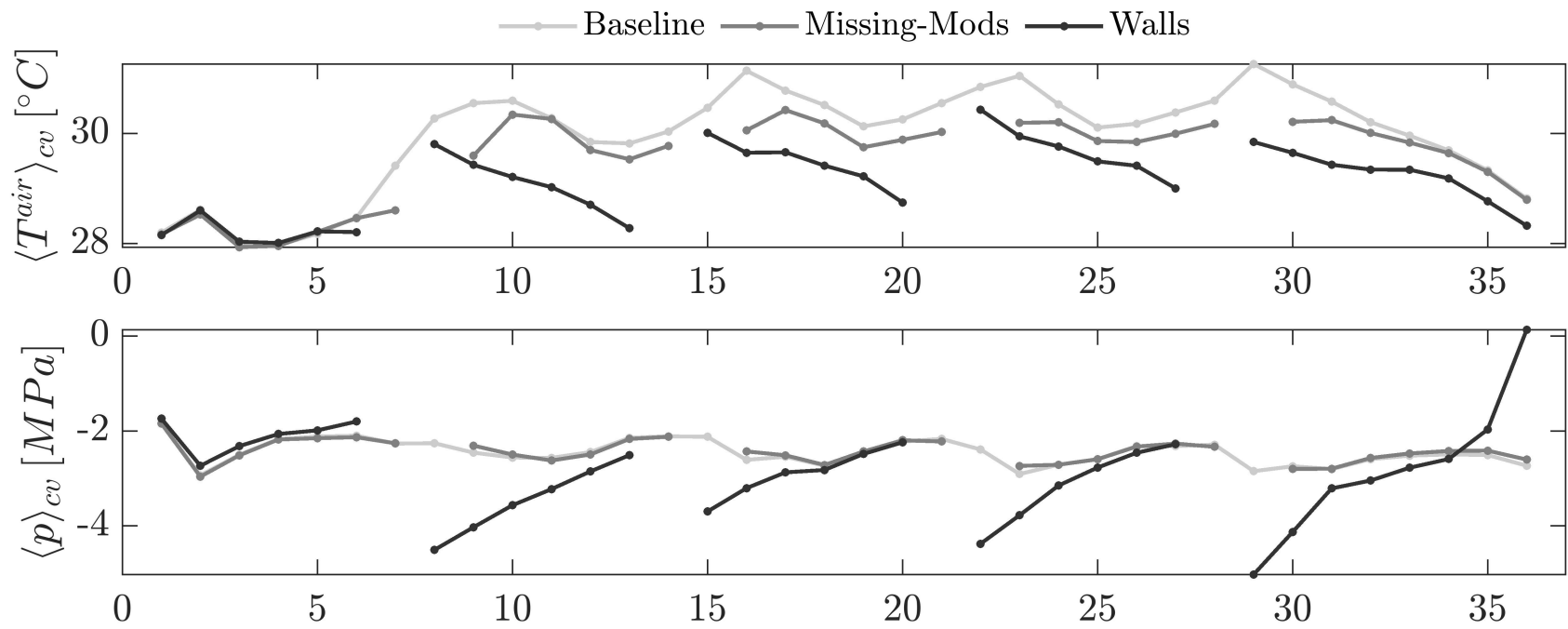
Thank you

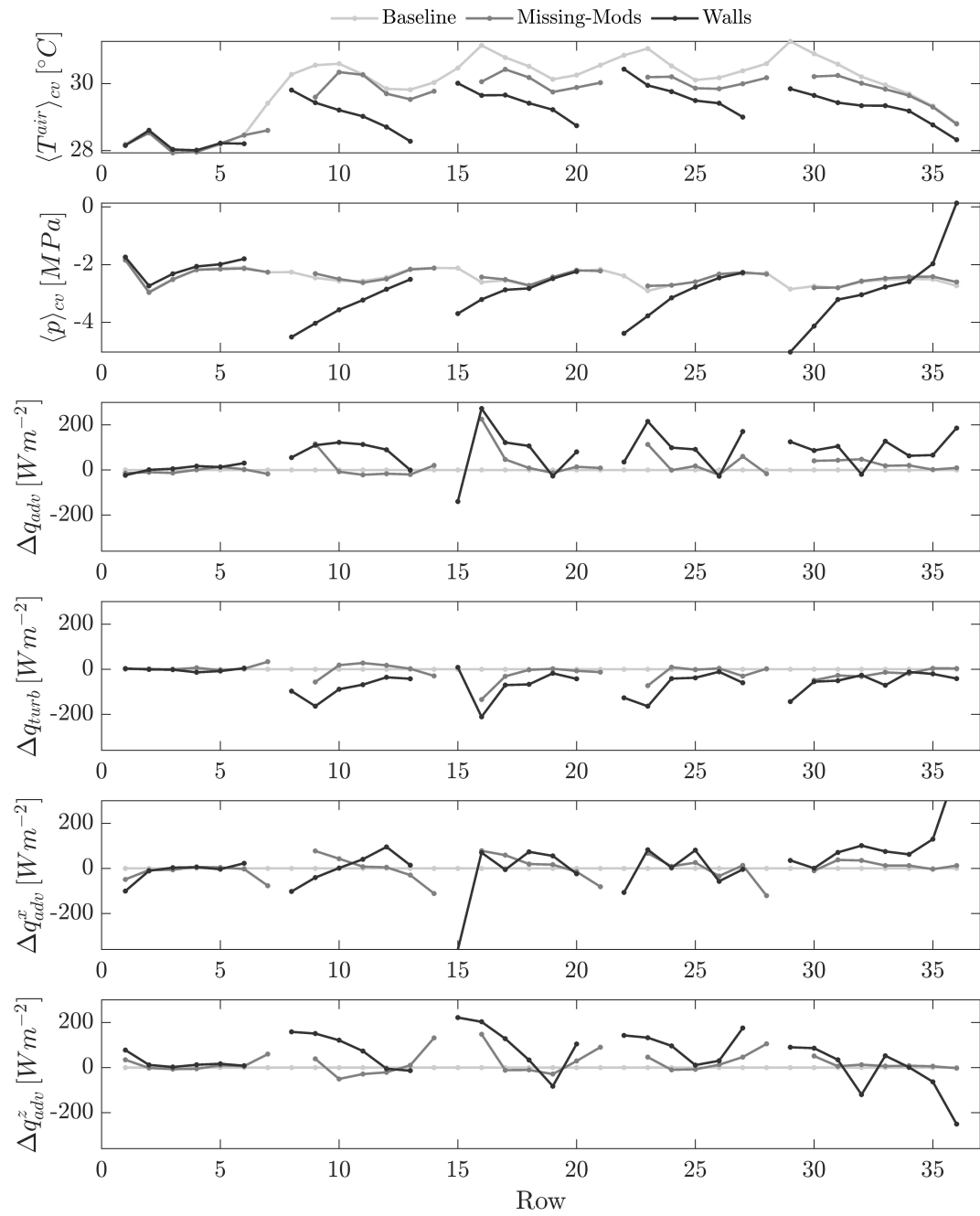
Brooke Stanislawski
brooke.stanislawski@nrel.gov

Barrier walls boost advective heat flux



$$\underbrace{\int_{A_k} [\bar{\mathbf{u}}_{\perp} (\bar{T} - T_{ref})] dA_k}_{q_{adv}}$$





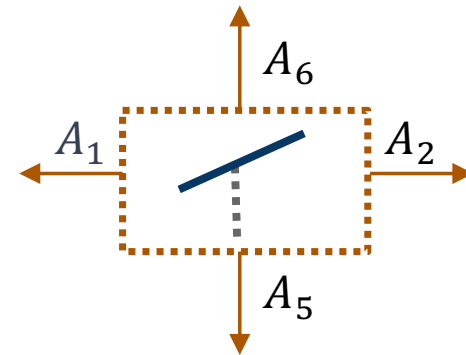
LES results are analyzed to compute the convective heat transfer coefficient h

Newton's Law of Cooling

$$h = \frac{q_m}{T_m - T_\infty} \quad [\text{W/m}^2\text{K}]$$

To solve for q_m , we start with the enthalpy equation

$$\rho C_p \frac{\partial \bar{T}}{\partial t} + \rho C_p \frac{\partial (\bar{u}_j \bar{T})}{\partial x_j} = -\rho C_p \frac{\partial (\overline{u'_j T'})}{\partial x_j} + \bar{s} \quad [\text{W m}^{-3}]$$



Integrating over the control volume

$$\underbrace{\rho C_p \int_{cv} \frac{\partial \bar{T}}{\partial t} dV}_{Q_{st}} + \underbrace{\rho C_p \int_{cv} \frac{\partial (\bar{u}_j \bar{T})}{\partial x_j} dV}_{Q_{adv}} + \underbrace{\rho C_p \int_{cv} \frac{\partial (\overline{u'_j T'})}{\partial x_j} dV}_{Q_{turb}} = \underbrace{\int_{cv} \bar{s} dV}_{Q_s [W]}, \quad [W]$$

$$Q_s = Q_{cv} = Q_m$$

Neglecting storage and converting volume integrals to surface integrals

$$q_m = \rho C_p \left[\underbrace{\sum_{k=1}^6 \int_{A_k} [\bar{\mathbf{u}}_\perp (\bar{T} - T_{ref})] dA_k}_{q_{adv}} + \underbrace{\sum_{k=1}^6 \int_{A_k} (\overline{\mathbf{u}'_\perp T'}) dA_k}_{q_{turb}} \right] \cdot [\text{W m}^{-2}]$$

where $T_{ref} = \langle T_{air} \rangle_{cv}$

MPMICE Equations

Conservation of mass

$$\frac{1}{V} \frac{D_r M_r}{Dt} = \sum_{s=1}^N \Gamma_{rs}$$

Conservation of momentum

$$\frac{1}{V} \frac{D_r (M_r \mathbf{u}_r)}{Dt} = \theta_r \nabla \cdot \boldsymbol{\sigma} + \nabla \cdot \theta_r (\boldsymbol{\sigma}_r - \boldsymbol{\sigma}) + \rho_r \mathbf{g} + \sum_{s=1}^N \mathbf{f}_{rs} + \sum_{s=1}^N \mathbf{u}_{rs}^+ \Gamma_{rs} \text{ +turbulence}$$

acceleration due to mean mixture stress
acceleration due to deviation of material stress from mean mixture stress
gravitational body force

model for momentum exchange among materials

rate of mass conversion among materials

Conservation of energy

$$\frac{1}{V} \frac{D_r (M_r e_r)}{Dt} = -\rho_r p \frac{D_r v_r}{Dt} + \theta_r \boldsymbol{\tau}_r : \nabla \mathbf{u}_r - \nabla \cdot \mathbf{j}_r + \sum_{s=1}^N q_{rs} + \sum_{s=1}^N h_{rs}^+ \Gamma_{rs}$$

model for heat energy exchange among materials

$$q_{12} = H_{12} \theta_1 \theta_2 (T_2 - T_1)$$

H_{12} is analogous to a convective heat transfer coefficient

closure relation for material stress

$$\boldsymbol{\sigma}_r = -p \mathbf{I} + \boldsymbol{\tau}_r$$